

Keck Next-Generation AO: Galactic Science

Michael Liu (chair)
Andrea Ghez
Tom Greene
Lynne Hillenbrand
Jessica Lu
Bruce Macintosh
Stan Metchev
Nevin Weinberg

CfAO workshop, March 2006

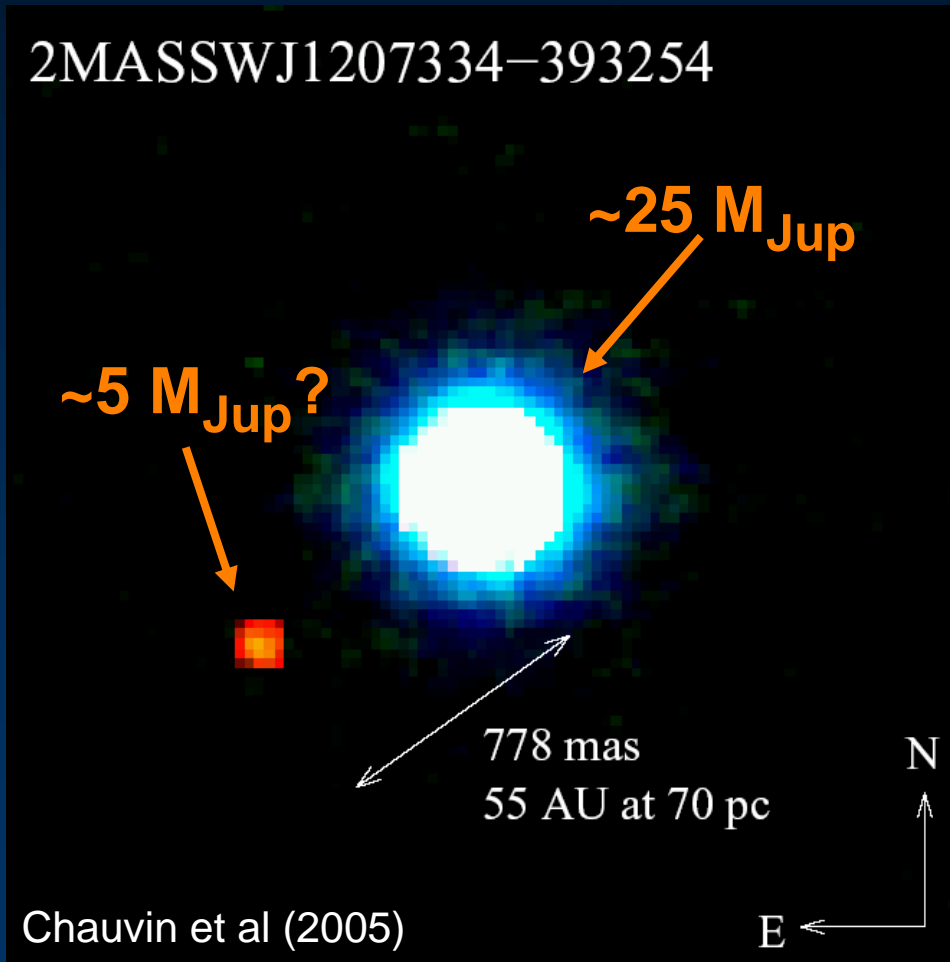
Key Science Cases

1. Planets around VLM stars & brown dwarfs
(Liu, Macintosh)
2. Galactic Center
(Ghez, Lu, Weinberg)
3. Debris Disks
(Metchev, Liu)
4. Circumstellar environments of protostars
(Hillenbrand, Greene)

Galactic Science: Main Requirements

- Narrow-field, high Strehl only.
- Optical and IR coverage needed.
- Some specialized capabilities needed
 - High contrast coronagraphy
 - Polarimetry

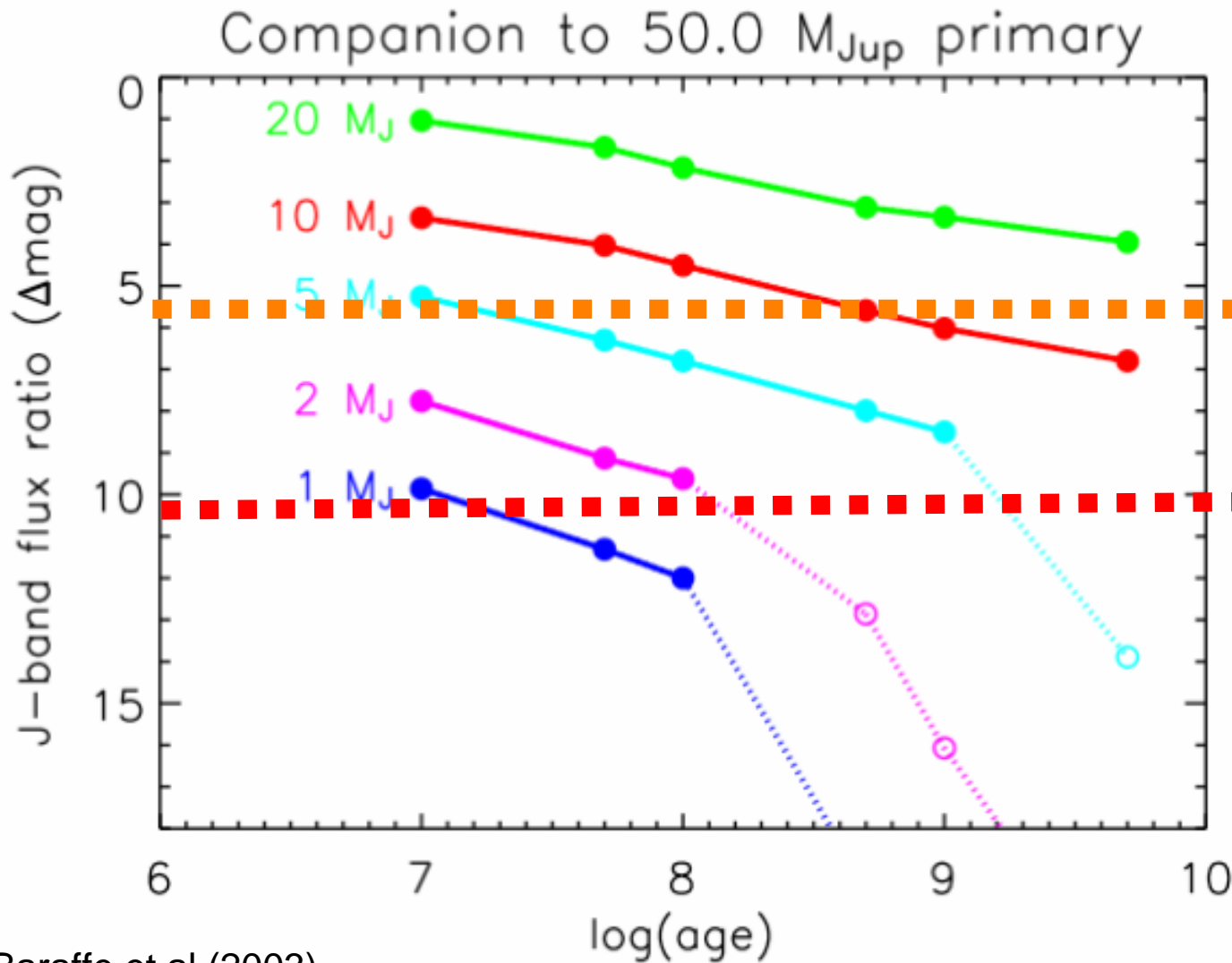
Planets around VLM stars & brown dwarfs



- Direct imaging of planets easier around VLM stars and brown dwarfs.
- Study objects with **SEDs** & **atmospheres** similar to “regular” exoplanets (but with different origin).
- High contrast LGS + IR tip tilt + high efficiency needed.

Contrast Requirements

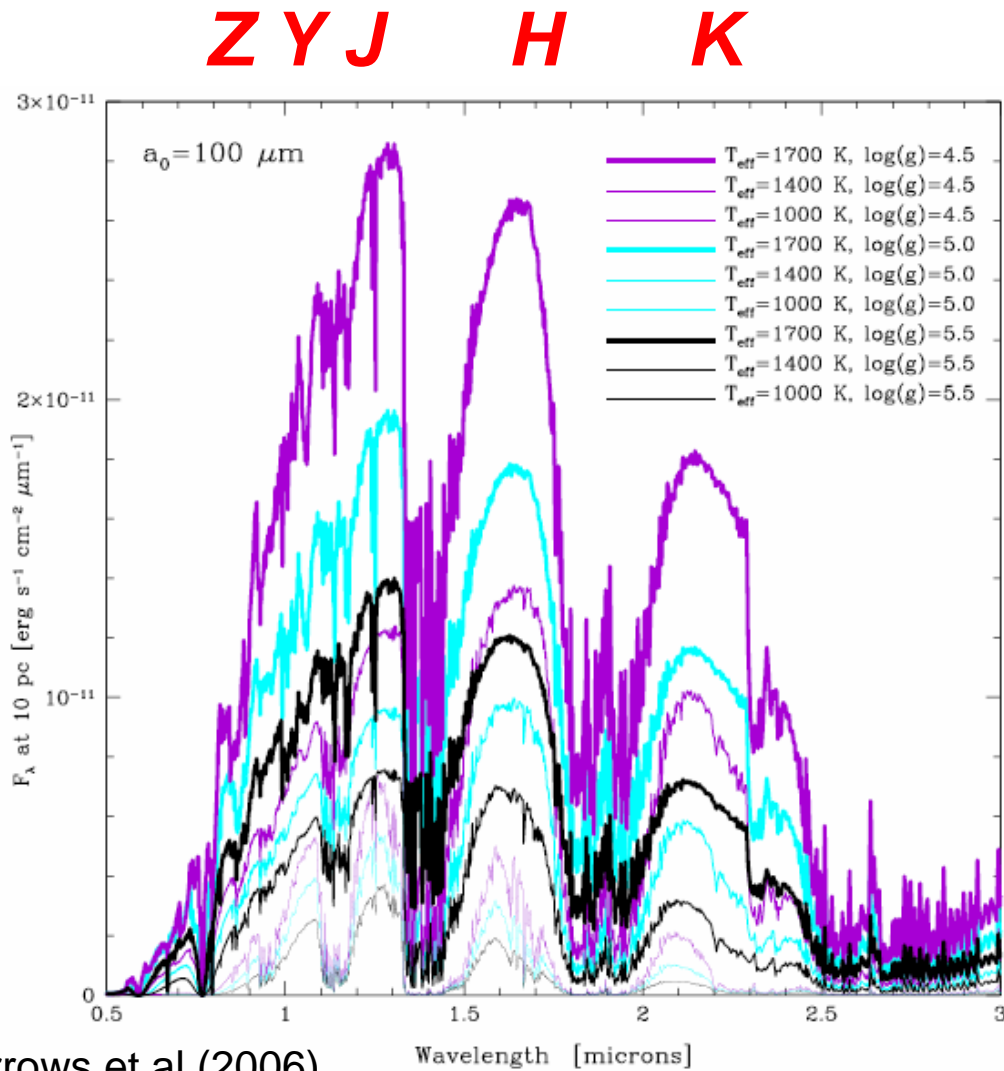
(known BD binaries are nearly all $< \sim 0.5''$)



LGS now
($r \sim 0.3''$)

KPAO
goal

Planet Characterization



Burrows et al (2006)

- **Diagnostics**
 - Surface gravity
Proxy for age/mass
 - Metallicity
 - Clouds

- **Observations**
 - Opt/IR colors
ZYJHKL' bands
 - Spectra
R ~ few 100

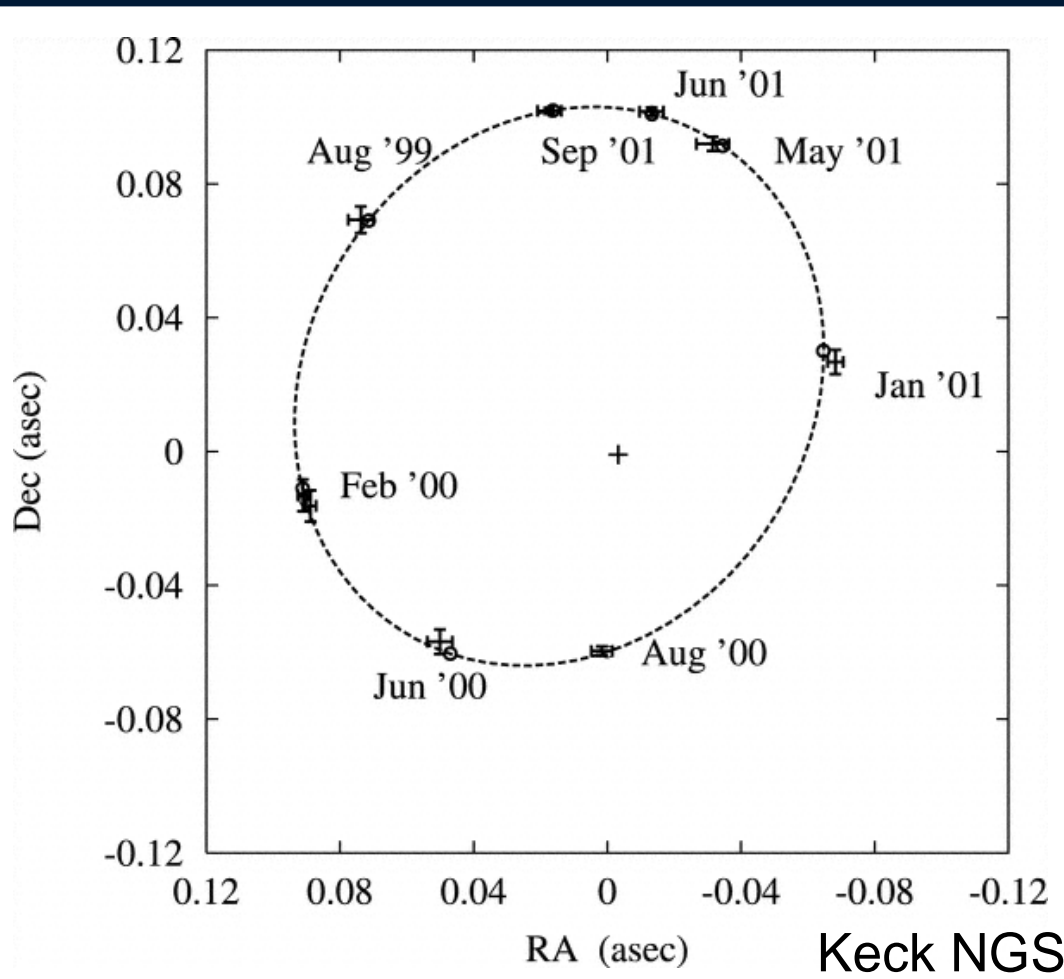
No shortage of science targets

1. Mostly all off-axis T/T stars (maybe)
 2. In general, tradeoff between age & distance...
- **Solar neighborhood**: $t \sim \text{Gyrs}$, $d < 20 \text{ pc}$
 - 2MASS/SDSS surveys of L & T dwarfs: $N \sim 500$
 - New IR surveys for Y dwarfs (**UKIDSS**, **VISTA**): $N \sim \text{dozens?}$
 - Volume-limited surveys (**Pan-STARRS**): $N \sim 1000\text{s}$
 - **Young field objects**: $t \sim 10\text{-}100 \text{ Myr}$, $d < 50 \text{ pc}$
 - Ongoing kinematic & spectroscopic searches: $N \sim 100\text{'s}$
 - **Star-forming regions**: $t < 10 \text{ Myr}$, $d \sim 150 \text{ pc}$
 - Ongoing wide-field IR surveys: $N \sim 100\text{'s}$

Simulation Plan

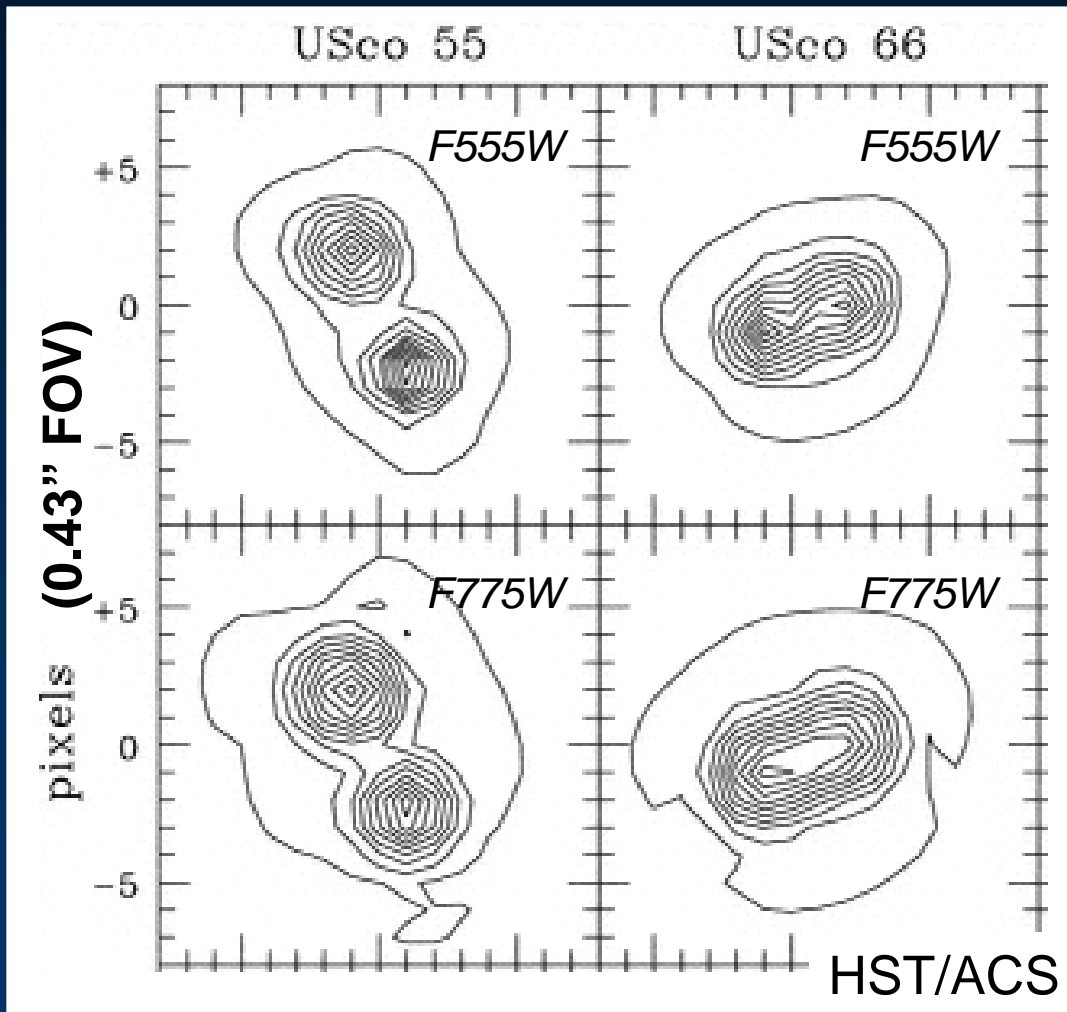
1. Estimate **pt source detection curve** for YJHK (*Δmag vs separation*).
2. Combine with substellar evolutionary models and notional target list (distance vs age).
3. Identify feasible/optimal bands for search & characterization.

Related low-mass science



- Dynamical masses of BD binaries.
Only ~2 systems have been measured so far.

Related low-mass science

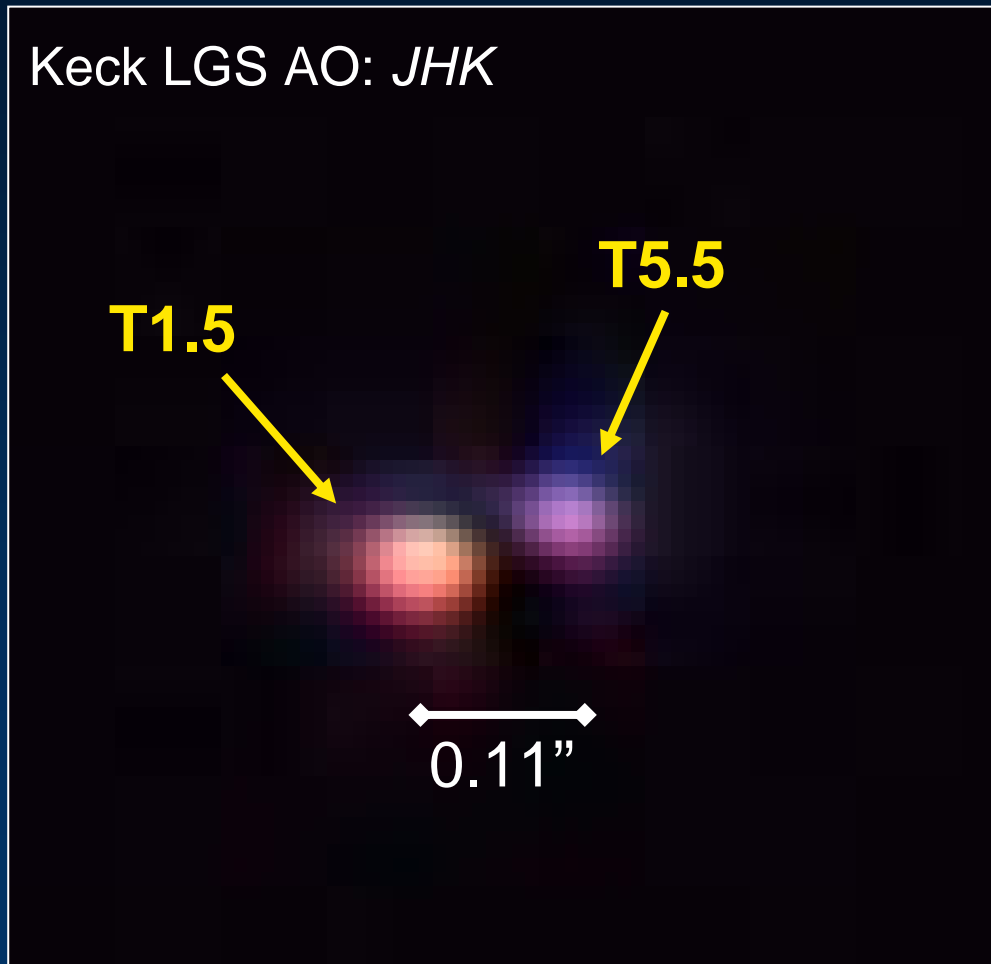


Kraus et al (2005)

M. Liu (IfA/Hawaii)

- Dynamical masses of BD binaries.
Only ~2 systems have been measured so far.
- Multiplicity of newly formed objects.
Testing BD formation models at young ages.

Related low-mass science



Liu et al (2006)

- Dynamical masses of BD binaries.
Only ~2 systems have been measured so far.
- Multiplicity of newly formed objects.
Testing BD formation models at young ages.
- Benchmark binaries.
Probing the physics of ultracool atmospheres.